The Trigonometrical Survey of India, (Communicated by Major J. T. WALKER.)

The following is the first of a Series of papers on matters of general interest connected with the Trigonometrical Survey of India, which it is proposed to extract from the manuscript volumes of the Survey, for publication in the Journal of the Asiatie Society. It is taken from the Introduction to the General Report of the North-East Longitudinal Series of triangles (G. T. Survey, Vol. XV.) drawn up under the Superintendence of Col. Sir Andrew Waugh, when Surveyor General of India, by J. B. N. Hennessey, Esq., 1st Assistant G. T. Survey.

The North-East Longitudinal Series derives its name from the circumstance of its following the course of the corresponding boundary of British India. It extends from the valley of the Dehra Dhoon to Purneah, connecting the northern extremities of the Calcutta Meridional Series and the celebrated Great Arc, measured by Cols. Lambton and Everest, on the meridian of Cape Comorin. Its object was to form the most direct connexion practicable between two base lines of verification, one measured in Dehra Dhoon, the other in Purneah. Thus it serves to close and verify the Meridional Series, 10 in number, which lie between the Great Arc and Calcutta Meridional Series and emanate from the longitudinal triangulation, connecting the Calcutta base with the Seronj base on the Great Arc in Central India.

This is the general system followed in the triangulation of India, which thus resembles in outline the form of a gridiron. At each angle of the gridiron, a base line is measured. The outer series form the frame-work on which the inner ones depend, and are especially valuable for the data they contribute towards the determination of the great problem of geodesy, the accurate measurement of the figure of the earth. By restricting the meridional, or inner series, to distances of 60 to 100 miles apart, all the necessary data for topographical operations are obtained, at a moicty of the cost that would be incurred in throwing a net work of triangles over the whole of India after the manner of European surveys, which require greater detail than is necessary in this country.

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The North East Longitudinal Series was originally intended by Col. now Sir George Everest, C. B. to have been carried along the mountains on the British frontier. But this design was abandoned in consequence of the refusal of the Nepalese Government to allow the operations to enter their territories. Consequently, after crossing the hills of Kumaon and Gurhwal, the triangles were brought down into the Terai near Bareilly, from which point they lie almost continuously in the marshy and deadly tracts which fringe the Himalaya mountains. Here Lt. Reginald Walker, a very able and promising young officer, fell a victim to jungle fever. Being alone and without medical assistance, he strove to reach Darjeeling, but was found dead in his dhooly, on its arrival at that station. Of the native subordinates, a large percentage, one year no less than a fourth, died of jungle fever. Sickness was frequent and severe. On more than one occasion a whole party had to be literally carried into the nearest station for medical assistance. The completion of the major, and more difficult portion of the triangulation is due to the ability, courage and perseverance displayed by Mr. George Logan, who died three years afterwards, from disease first contracted in the Terai during these operations.

Owing to the proximity of the triangulation to the mountain ranges, the whole of the chief peaks were seen from the principal trigonometrical stations, and fixed by measurements with the first class instruments employed for the mutual observations between the stations themselves. These are called the "Principal Observations," for on them, the accuracy and value of the series, as a whole, depend. They are therefore taken with the largest and most powerful theodolites, which are expressly constructed for the Indian Survey, and furnished with micrometer microscopes, instead of verniers, for reading the graduations.

The employment of such instruments in secondary operations has the advantage of enabling the observer to attain as great accuracy by a few observations as by many with second class instruments; thus time is saved and reliable measurements of the higher mountains can be taken during the short intervals when their usually cloud-capped summits are unfurled to view.

The following extracts are chiefly relative to the computations for determining the heights and positions of the principal mountains.

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A table of the resulting elements is given, together with a memorandum specifying the mountains which could be identified as having been previously observed by other surveyors. J. T. W.

Of the Secondary Mountain Triangulation.

57. The magnitude of the triangles for determining the positions of the hill peaks, and other unavoidable peculiarities attendant on the operations in general, have necessitated some few departures from ordinary precedents in the performance of the required calculations. These may be briefly noticed.

Identification .--- The primary difficulty which the computer 58. meets with is, in the identification of the numerous points whose positions have been determined. Observed by different persons, after long intervals or from different points of view under the disadvantages of altered aspects, the same hill will be found noted in the angle books under various characteristics. For instance, Mont Everest was called v by Colonel Waugh, n by Mr. Nicolson and b by Mr. Armstrong, while the peak XXXVIII. is named n^2 at one station of observation, n^3 at another and "I west peak" at a third, by the same observer. This plurality of characteristics, under the circumstances, is clearly unavoidable. It remains to state how the required identification was effected. The principal series was first carefully projected on a scale of 4 miles to the inch, and the several rays emanating from stations of observation were next exactly drawn. The intersection of these rays, assisted by the characteristics forthcoming in the angle books, more or less distinctly defined the points sought for. This was treated as an approximate identification, whereby the bases required from the principal series and expermiental triangles to be computed became known. The former were then obtained in the ordinary way, by means of the contained angle and logfeet of the including sides, for which computation the following well known formula was found useful,

$$\tan \frac{1}{2} (A - B) = \tan (45 - Q) \operatorname{Cot} \frac{C}{2}$$

wherein $\tan Q = \frac{b}{a}$

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With the bases so found, the triangles were, as implied, first experimentally computed, an accordance of the numerous common sides demonstrating an identity of the several characteristic letters. In those cases where any want of demonstration existed, the point was rejected.

59. Such identification imposes no *experimental* calculation when the points observed are clearly isolated from each other. For instance XI. or Jannoo, XIII. and Mont Everest or XV. were readily identified by the angular projection. But as in the cases of XLIII., XLIV. and XLV. it is evident that nothing short of actual computation will separate the points in the group. The numerous experimental triangles by which non-identity was proved, as also the triangles for bases are not shown in this volume. The last mentioned triangles were about 450 in number, and the former also involved considerable labour.

60. Spheroidal excess.—The two formulæ for spheroidal excess, viz., that involving two sides and the contained angle, and the other in terms of the base and the three angles, were respectively employed in the triangles for bases and in those to Himalayan points. In the latter case however, the spherical angle opposite the base c could, in the first instance, be only roughly found from the equation $\pi - (A + B) = C$, wherein A and B are spherical angles. Whence C was taken too small by the whole spheroidal excess. Now, as this latter frequently exceeds 100 seconds, it was sometimes required to find the excess approximately, next to correct the angle C, and then with this value of C, to recompute the excess finally. In other respects the Triangles were calculated as usually done.

61. Synopsis of sides.—The values of the sides in feet thus obtained were recorded in the form of a synopsis, and this paper was completed by finding the logarithm to the mean of these values, as well as the miles corresponding to the same.

62. Latitude and Longitude.—The computer was now prepared to deduce the required latitudes and longitudes, which was done in this wise. With the latitude and longitude of any station of observation A, the aximuth thereat of point n, and the mean distance from the synopsis of sides A to n, the latitude and longitude of n from A were found. Similarly values of latitude and longitude were obtained from the other stations of observation, and a mean of all these values was taken as *the* latitude and longitude of n. 63. The computation of heights was performed in the usual manner, until the estimation of terrestrial refraction was arrived at. The process adopted for this purpose may be briefly stated thus.

64. Estimation of Terrestrial Refraction.-If the contained arc

be represented by c, and terrestrial refraction by r, then -=f

the factor, or "decimals of contained arc." Whereby if f be given, then r = c.f may be computed. From want of a more accurate method of determination, it is usual to adopt that mean value of f, for finding the height of an inaccessible point, which may be forthcoming from the reciprocal observations at visited stations. For instance if A, B, C, D, be points of the last mentioned order, then



in the ordinary course of computation, there will result three values of f at A, as many at C, and two values each at B and D. The mean value of f at each station would therefore be adopted in computing the height of an inaccessible point H. To take a real case (at random). The values of f at Batwya T. S. (1) are + 0.011, -0.017, + 0.065 and + 0.013. Wherein the greatest difference is no less than .082 of the contained arc. On the other hand, the values of f at hill stations of observation, will always be

found accordant within far smaller limits.

65. The conclusion drawn from the foregoing is evidently this. That at plain stations, and when the object observed is placed on an ordinary tower, the value of f determined from any given ray A B, is not necessarily applicable to any other ray A C. Whereas all rays of light at hill stations from terrestrial points appear to be nearly equally refracted. These phenomena are clearly traceable to local causes.

66. But of the two mean values of f, one obtained at a mountain station of observation, and another deduced in the plains, it is evident that the former is more trustworthy, and hence it appeared desirable, that the latter should be obtained in terms of the former.



67. Process of estimating terrestrial refraction.—Let A, B, C, D, (vide figure) be plain stations, T and S stations on the Sub-Himalayas, and I. to IV. inaccessible points on the range of perpetual snows. Let the values of f at T and S equal respectively f and f_s . We may deduce from these, two trustworthy values of the heights of I. and II. Calling this mean height of $I = I_m$, and remembering that we have elevation (E) at C of I, as also the contained arc for C I = (c) given, it is clear that the values of f at C, corresponding to I_m may be found. Let this value = f. Proceeding in the same manner we shall find $f_c = \frac{f^{\tau} + f^2 + \ldots + f_n}{n}$.

Similarly $f_{\rm D}$ &c., may be obtained, and with f_c , $f_{\rm D}$ &c., may be computed III_m , IV_m &c., from which again in turn may be found the values of f for the other plain stations from which III, IV &c, have been observed. By this process the computed values of f are determined nearly in terms of f_t and f_s , errors of observation not being taken into account. It remains to mention how f_s and f_t were obtained.

68. The computations originate from Senchal and Tonglo hill stations, at which stations, the following mean value of f was in the

first instance adopted. The selection has been made to the exclusion of those values obtained from short sides.

Deduction.—	-Doom Dangi Senehal	$\Big\} f = .07617.$
	Thakoorganj Senehal	$\Big\} f = .07636.$
	Doom Dangi Tonglo	f = .07915.
	Thakoorganj Tonglo	f = .07849.
	Senchal Tonglo	$\Big\} f = .06201.$
	Tonglo Darjeeling	$\Big\} f = .08043.$
	Mean	$f = .0744. = \frac{1}{13 \cdot 2}$ nearly.

69. With this value of f, the heights from Senchal and Tonglo were computed, and the mean of these values, as also the differences between each value and its mean, were next found. The heights were now corrected, in such wise, that when the heights deduced from Senchal are compared with the mean heights already mentioned, the greatest + and - differences should be numerically equal. The same process being gone through at Tonglo, H. S., there resulted the mean values of f, which have been employed for that station and for Senchal. These values will be found recorded in the heights herein given, and it will also be found, that they have been employed for all heights of the *Sub*-Himalayas observed at Senchal and Tonglo hill stations.

70. It may be useful to remember, that if there be two points A and B observed from O, whose heights respectively are h_a and h_b determined by a certain value of f at $O = f_o$. Also if d_a equal corrected geodotic distance O to A, and $d_b = O$ B. Then if f_o vary, so that h_a (the height of A computed from O) changes by $\pm \delta_a$, and h_b by $\pm \delta_b$, so will $\pm \frac{\delta_b}{\delta_a} \propto \frac{d_b^2}{d_{a2}}$. Hence should the foregoing method for finding the value of f at plain stations in terms of the observed value at hill stations, be hereafter ever adopted, it will be found advantageous to construct a table of the squares of the distances in miles, for this purpose.

71. The general principle of procedure is now apparent. But as

will be remarked, the process described is only applicable so long as a continuous connection is preserved, between the stations of observation and the points observed. In the observations under consideration, there occurs a blank space between points LII. and LIII whence the method described was no longer applicable beyond the former point. But it fortunately happens that LIII. and succeeding points are observed from hill stations, whereat, as already mentioned, the values of f are liable to but triffing variation. The mean value of f in these cases was deduced in the ordinary way as mentioned at para. 64. The following is an example of this method.

At Jagesar, H. S. the values of

$$(f) \text{ are } \begin{cases} .04485. \\ .04528. \\ .04876. \end{cases}$$

Mean f adopted at Jagesar, H. S. .04630.

72. Values of f tabulated.—The values of f employed in these calculations may be tabulated thus.

Height	N	4-1.			0	Denominator		
above sea	in ames of S		\mathcal{J} .	vulgar frac-				
level.			tion.					
Feet.								
8610	Senchal, H. S.				.0815	12.2657		
319	Doom Dangi, T. S.				.0744	13.4374		
7169	Darjeeling, H. S.				.0885	11.2945		
6884	Birch Hill, S.				.0864	11.5737		
273	Thakoorganj, T. S.				.0775	12.9066		
10084	Tonglo, H. S.				.0711	14.0550		
251	Banderjoola, T. S.	•••			.0811	12.3317		
237	Menai, T. S.				.0753	13.2852		
242	Baisi, T. S.				.0743	13.4677		
226	Harpoor, T. S.	•••			.0727	13.7637		
242	Ladnia, T. S.				.0746	13.4025		
263	Janjpati, T. S.				.0731	13.6705		
254	Mirzapoor, T. S.				.0736	13.5775		
231	Jirol, T. S.	•••			.0735	13.6008		
282	Sinereah, T. S.		•••		.0753	13.2797		
268	Boolakipoor, T. S.	• • •	•••		.0728	13.7429		
259	Batwya, T. S.				.0714	14.0093		
320	Torharwa, T. S.				.0847	11.8002		
357 .	Morairi, T. S.		•••		.0791	12.6429		
353	Soopoor, T. S.	•••			.0813	12.3031		
355	Banarsi, T. S.	•••	•••		.0937	10.6681		
344	Saonbarsa, T. S.				.0870	11.4928		
350) Bharmi, T. S.				.0787	12.7054		
329	Poorena, T. S.				.0805	12.4154		
358	Ghaos, T. S.)	.0875	11.4292		

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Height above sea level.	Names of Sta	tions			<i>f</i> .	Denominator of vulgar frac- tion.
412	Toolsipoor. T. S.				.0763	13.1058
478	Anarkali, T. S.				.0744	13.4432
7732	Jagesar, H. S.				.0463	21.5983
6994	Birond, H. S.		•••		.0652	15.3374
10101	Khankra, H. S.	•••		•••	.0579	17.2652
8526	Soonchalia, H. S.	•••			.0624	16.0256
6946	Ghoongti, II. S.	• • •	•••	••••	.0652	15.3374
7079	Ranigarh, H. S.	•••		• • •	.0687	14.5624
5675	Mabegarh, H. S.	•••		•••	.0750	13.3333
7371	Ghandial, H. S.	• • •			.0698	14.3266
12541	Kiderkanta, H. S.	•••			.0480	20.8377
9946	Nagtiba, H. S.	• • •	•••	•••	.0521	19.1902
2970	Dhoiwala, H. S.	•••	•••	•••	.0628	15.9363
7454	Banog, H. S.	•••	+ r s	•••	.0612	16.3479
3161	Amsot, H. S.	•••	•••		.0565	17.6897
11997	Chur, H. S.	•••	•••	••	.0530	18.8857

73. Conclusion deduced from foregoing table.—Now since $Sin \leq incidence$

----- = 1 + m in the mean state of atmosphere and at Sin \angle refraction

the level of the sea, and also, since the quantity m varies with the density of the atmosphere, so that when the density of the air is only the nth part of what it is at the level of the sea, the refractive power is

there only $1 + \frac{m}{n}$, it might have been expected from these tabulated n

results that in the first instance, $f \alpha \frac{1}{\text{height of station of observation}}$

No such law, however, is to be found unless the numerous exceptional cases be excluded to make a rule.

74. Wherefore it appears, that the law of variation in f due to variation in the density of the atmosphere, consequent on variation in height, is completely absorbed and lost sight of in the irregular variations, arising from local causes and also from the unavoidable imperfections of observation to points so ill-defined as the apices of snowy mountains.

75. Finally it is to be noticed that the foregoing method is acknowledged to be imperfect and unsatisfactory, but compared with the ordinary mode of finding f from reciprocal vertical observations,

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it is believed that the values herein determined are a nearer approximation to the truth.

76. Notices certain refinements not appreciable in these operations.—In concluding the remarks on these computations, it may be interesting to notice certain refinements in calculation which have not been deemed applicable to these operations. For instance, the spheroidal excess and the contained arc might have been computed by more rigorous processes, but that the refinement would have been purely of an arithmetical nature. Again the formula for latitude and longitude has not been employed beyond its fourth term, because the remaining terms are difficult of arithmetical expression and would besides have given no results commensurate with the labour necessary to compute them. Similarly the chord correction is neglected in these heights, amounting as it does in the extreme case of Menai to Mont Everest, or XV, to no more than a foot.

77. There remains to notice one other correction also herein not taken into account, of which it may be remarked, that, under existing circumstances it would partially cancel the chord correction, if both these refinements were introduced. This correction may be stated thus.

78. Ordinarily, in the formula for computing difference of height, it is sufficiently accurate to assume the given arc (or distance) to belong to a circle, whereas in reality, it is a portion of an ellipse. If the correction due to this assumption = x b, then it can be shown that $x b = (\nu a - \cos \lambda_b \ K) - (\nu_b - \cos \lambda_a \ K)$, wherein K $= \left\{ \nu_b \sin \lambda_b - \nu_a \sin \lambda_a + \frac{N}{M} \left[(M + \nu_a \cos \lambda_a) (M - \nu_a \cos \lambda_a) \right]^{\frac{1}{3}} - \frac{N}{M} \left[(M + \nu_a \ \cos \lambda_b) (M - \nu_b \ \cos \lambda_b) \right]^{\frac{1}{2}} \right\} \text{Cosec } \delta \lambda.$

It is sufficient to remark in this place, that in the extreme case of Menai, T. S. to Mont Everest or XV. the correction x b = only 0.3 of a foot.

79. Magnitude of these operations illustrated.—Lastly it may be interesting to notice, that the area of the largest triangle to points on the Himalaya mountains (No. 297) is about 1706 square miles, its spheroidal excess being 106". The longest side, Anarkali, T. S. to XXXIX. is equal to 151 miles, and its corresponding contained arc

is 7886'' = about the $\frac{1}{164}$ th part of a circle described around our planet. And if the principal and mountain operations of the North East longitudinal series be taken together, they will be found to cover somewhat more than the $\frac{1}{3182}$ portion of the entire earth's surface; or, taking the land at half the expanse of water, about 1061 such series would cover every portion of the former.

80. Accuracy discussed.—And with regard to the accuracy of the mountain results, it is evident that the same estimate cannot equally apply to a peak with a sharp conical apex, and to a mountain whose summit represents a saddle back or an even bluff. Prominent amongst the accurately determined points are XIII. Mont Everest or XV. and XLII. or Dhoulagiri, both in respect to geographical position and height above sea level, but though such points are far more numerous than those which exhibit comparatively large differences between the several values composing their mean results, yet it is suggested that the synopsis of latitudes and longitudes and the paper of heights should be consulted before adopting a point, if necessary for rigorous purposes.

81. The same estimated.—It is estimated, that on an average, the points on the Himalaya mountains are correct in latitude to $\frac{1}{4}$ of a second and in longitude to about $\frac{1}{2}$ that quantity. The heights are probably true to 10 feet, but this last estimate must be qualified by the consideration that they are all too low from the deflection due to mountain attraction.

82. Why mountain attraction was not determined.—In the original design of these operations, it was intended that the deflections in azimuth and in the meridian due to the attraction of the Himalaya mountains should be estimated along the principal series by suitable celestial observations, but this intention was relinquished owing to the considerable delay it entailed.

84. Area and cost.—The area covered by these principal and secondary operations amounts to about 61,815 square miles. But the piecemeal nature of work, the long intervals which frequently occur, and the unavoidable employment of the North East longitudinal series partly on other duties, make it a difficult and unsatisfactory process to attempt finding the cost of these operations. As an approximation, however, it may be stated that this cost does not exceed Rupees 2 per square mile, $1862.]^{\circ}$

Table of characteristic marks, for the snowy peaks of the North East longitudinal series, great Trigonometrical Survey of India, and identification with other authorities.

Final Numeral and Name adopted.	Country.	Identification with other authorities.
I. or Choomlari, II. or Gipmochi, III. or Porohoonri, IV. or Choomoonko, V. or Black rock,	Tibet. Bhotan. Tibet & Sikkim. do. do.	Named by Dr. Hooker, Donkiah. Named by Dr. Campbell, Chola. Named by Dr. Campbell, Gnaream.
VI. or Narsing, VII. or Pandim, VIII. or Kanchinjinga, IX. or Kanchinjinga, X. or Kabroo, XI. or Jannoo,	Sikkim. do. do. Nepal & Sikkim. do. Nepal.	
XII	do.	
XIII	•do.	
XIV	do.	1
XV. or Mont Everest,	do.	•
XVI	do.	
XV11	do.	Colonel Crawford's A.
AVIII	do.	Colonel Crawford's B.
$\frac{\Lambda I \Lambda}{N N}$	do.	
AA	do.	Colonel Crawford's C.
AA1	do.	Colonel Crawford's D.
AAll	do.	Colonel Crawford's F.
XXIII	do.	
XXIV	do.	
AAV. or Dayabang,	do.	Colonel Crawford's L. or Dayabang.
$\Delta \Delta V I$	do.	υ.
AAVII	do.	
XAVIII	do,	
$AAIA. \dots \dots$	do.	
$\Delta \Delta \Delta$	do.	
AAAL	do.	
VVVIII	d0.	c)
VVVIV	d0.	
VXVV	do.	
XXXVI	do.	•
XXXVII	do.	
XXXVIII	do.	
XXXIX.	do.	
XL	do.	
XLI.	do.	· [airi]
XLII. or Dhoulagiri	do.	Capt. Webb's Dhawalagiri (Dhoula-
XLIII.	do.	The cost of Dilandia Biri, (Dilouia-
XLIV.	do.	
XLV.	do.	
XLVI	do.	· · · · · · · · · · · · · · · · · · ·
XLVII	do.	
XLVIII	do.	
XLIX	do.	

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Gurhwal.

Final Numeral and Name adopted.	Country.	Identification with other authorities.
L LI LII LIII. or Api, LIV. or Panchachuti, LV	Nepal. do. do. do. Kumaon. do.	Capt. Webb's XXIII. (Api.) Capt. Webb's XIX. Capt. Webb's XVIII.
LVI. or Nandakut, LVII LVIII. or Nandadebi, LIX.	do. do. Kumaon & Bri- tish Gurhwal	Capt. Webb's XV. Capt. Hodgson and Lt. Herbert's A. No. 2: Capt. Webb's XIV.
LXI. or (East) Trisool, LXI	British Gurhwal. Kumaon and Bri- tish Gurhwal.	Capt. Webb's XIII. (East) Trisool. Capt. Hodgson and Lt. Herbert's P or A. No. 3, Capt. Webb's N.
sool	British Gurhwal.	Capt. Hodgson and Lt. Herbert's A. No. 1; Capt. Webb's XII. or West Trisool.
LXIII,	do.	• · · Capt. Webb's XI. (Nandakna.)
LXVI LXVII. or Kamet or	do.	Capt. Webb's K.
Ibi Gamin,	Tibet and British Gurhwal.	Capt. R. Strachey's Kamet, named by* Messrs. Schlagintweit Ibi Gamin.
LXVIII. or Nilakanta, LXIX. or Badrinath,	British Gurhwal. do.	Capt. Webb's IX. (Nilakanta.) Capt. Hodgson and Lt. Herbert's B. Middle peak Badrinath, Capt. Webb's VIII.
LXX	do.	Capt. Webb's VI.
LXXI	do.	Capt. Webb's G.
LXXII. or Kedarnath,	Gurhwal and Bri- tish Gurhwal.	Capt. Hodgson and Lt. Herbert's D. or Kedarnath, Capt. Webb's III.
LAAIII	Cumberal	Mr. Keelan's a. Capt Hadreen and It Happart's M
gar,	5 Gurnwai.	or Mont Moira, Capt. Webb's 1.
LXXV. or Jaouli, LXXVI. or Bus Peak	do.	Capt. Hodgson and Lt. Herbert's C. or Jaouli, Mr. Keelan's <i>i</i> .
or Srikanta,	do.	Capt. Hodgson and Lt. Herbert's G. or Srikanta. Mr. Keelan's d. Mr. Mulheran's I. or Srikanta. Mr.
poonch,	do.	Capt. Hodgson and Lt. Herbert's Great E. or Banderpoonch. Mr. Keelan's a. Mr. Dyer's l.
LXXVIII	do.	Capt. Hodgson and Lt. Herbert's Low E.
LXXIX. or Sargoroen,	do.	Capt. Hodgson and Lt. Herbert's H. Left peak.

SERIES.	
VAL	
UDIN	
GITI	
LON	
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General Alphabetical List of Latitudes, Longitudes and Heights.

Remarks.		
District.	British Sikkim. Do. Darjeeling, British Sikkim. Landour Hills, N. of Dehra. Do. Do. Do.	Denra Duoon.
Heights above sea · levcl.	feet. 7169 7383 7485 7383 7485 7308 7620 6620	7310
Longitudes.	 , ', ' 88 18 36 88 18 32 88 18 32 88 18 32 88 59 22 81 58 52 81 58 52 82 0 45 83 56 79 58 12 79 58 12 79 58 12 79 58 12 79 58 23 78 6 58.71 78 6 58.71 78 7 37 78 7 37 	02.2 0 21
Latitudes.	 , ', ' 27 2 52 27 2 52 27 2 49.65 26 6 18 27 31 10 27 41 49 27 41 49 27 31 56 27 53 54 27 53 54 27 53 54 30 27 36 30 27 36 30 27 36 30 27 14 	21.76 VI US
Names of Places.	 Darjeeling Church, N. W. spire, Darjeeling, Campbell's (Dr.) centre chimney, Darjeeling, H. S., Kishanganj Rajah's Noubatkhana, Kishanganj Rajah's Noubatkhana, Bhinga Fort, 	Dehra Dhoon Ubservatory Station,
No.	1309(1 1310(1 1310(1 181 650 841 873 1194(1 1194(1 1326(1 1328(1 1328(1 1328(1 1328(1 1328(1 1328(1 1328(1 1328(1 1328(1 1328(1)(1))(1)(1)(1)(1)(1)(1)	1720

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		Remarks.	
ntinued.)		District.	Tibet Blotan. Tibet and Sikkim. Do. Do. Do. Do. Do. Nepal and Sikkim. Do. Do. Do. Do. Do. Do. Do. Do. Do. Do
ES-(Co	anges.	Heights above sea level.	fcet. 23944 14518 23944 14518 23186 17325 19146 27815 28156 24015 28156 24015 28156 24015 28156 282500 282500 282500 282500 282500 282500 2825000 282500
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The Latitude depends on the value of that element adopted for Kalianpoor Station $= 24^{\circ}$ 7' 11''.262,

The Longitude is referrible to the old value for the Madras Observatory $= 80^{\circ} 17' 21''$ to which a correction of -3' 25''.5 is applicable to reduce to the value adopted by the Admiralty and Royal Astronomical Society or -3' 1''.8 to reduce to the result of Taylor's observations up to 1845.

The Heights' originate from the mean sea level, observed in Kydd's Dock-yard, Calcutta.

The Trigonometrical Survey of India.